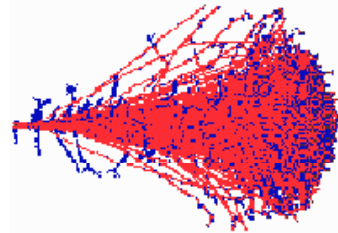


# Principles of Spectrometry

# Principles of Spectrometry

- Radiation detection requires that radiation interacts with detector material.
- Energy deposition (not all detectors...!).
- Charged particles **continuously** transfer their energy to the medium.

<http://www.srim.org>

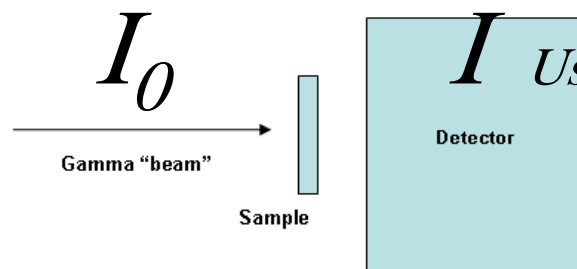


**Energy Loss:**  
Electronic or nuclear

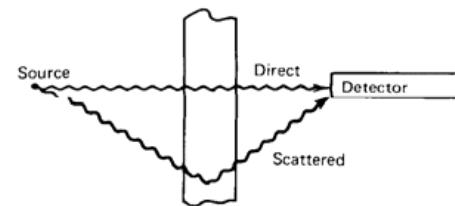
- Gammas and X-rays interact (if they do) **catastrophically**.

<http://physics.nist.gov/PhysRefData/Xcom/Text/XCOM.html>

**Attenuation:**  
Absorption or scattering



*Usually corresponds to net area under the full energy peak.*



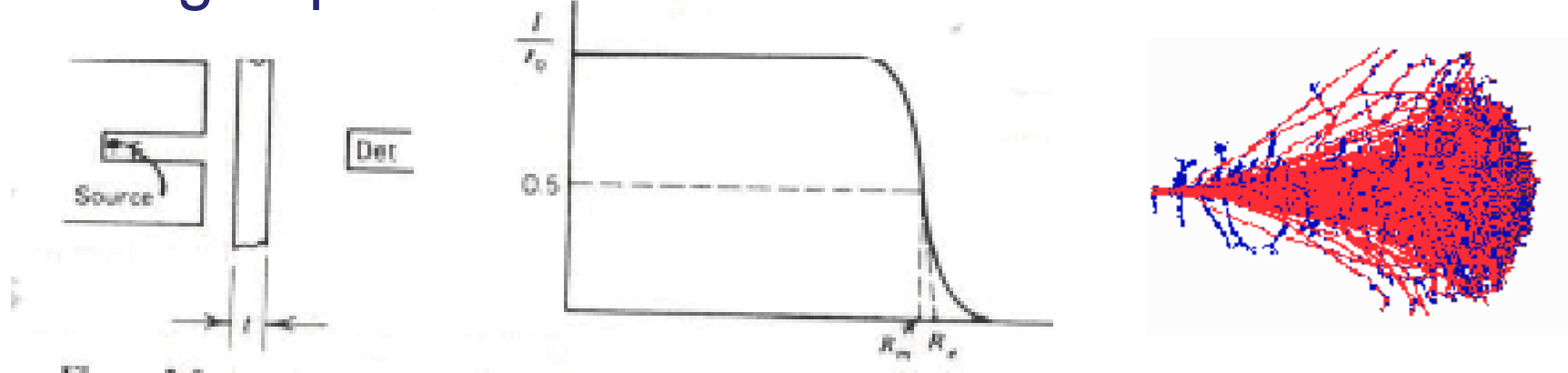
- Photon interactions produce **secondary electrons**.
- Neutron interactions produce **secondary heavy charged particles**.  
Could be secondary gamma (Shielding).

**If not,  
then  
build-up.  
Good or  
bad  
geometry.**

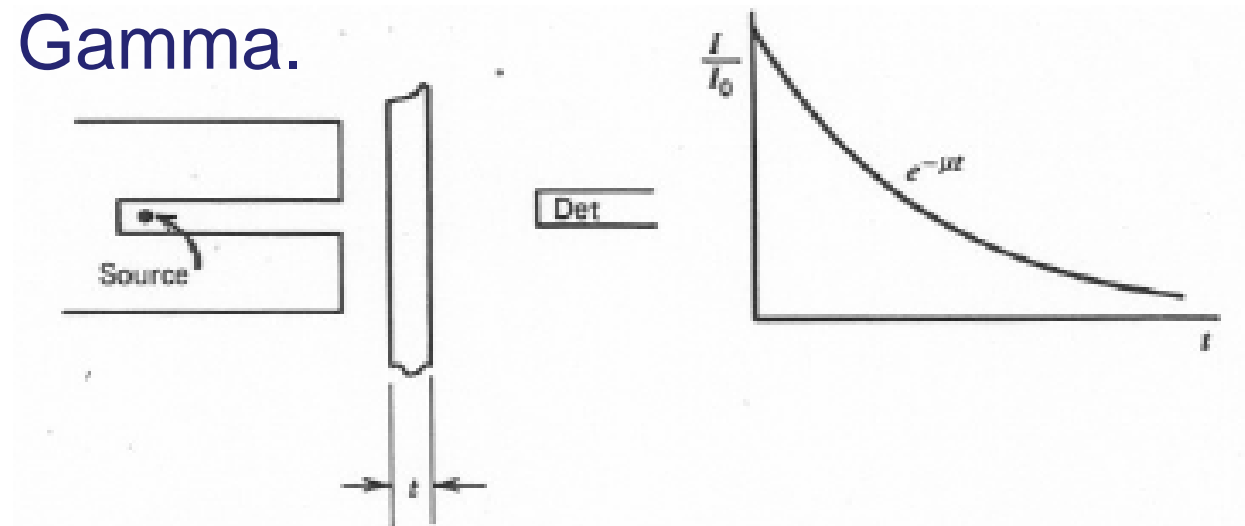
**Collimation.**

# Principles of Spectrometry

## Charged particles.



## Gamma.

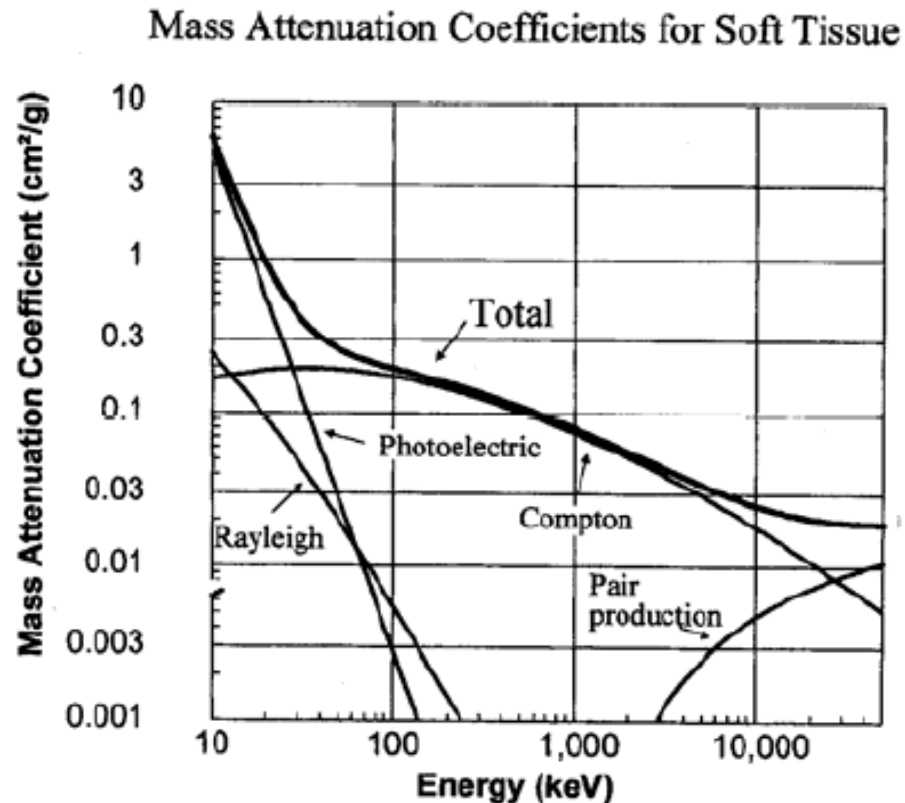


# Principles of Spectrometry

## Interaction of gammas

- Partial or complete transfer of photon energy to electrons.
- Many types of interactions, but three (**or four**) are important.

Buildup  
Factor



$$I = I_0 e^{-\mu x}$$

- **Attenuation coefficient and mass attenuation coefficient.**
- **Compare to macroscopic cross section  $\Sigma$  for neutrons.**

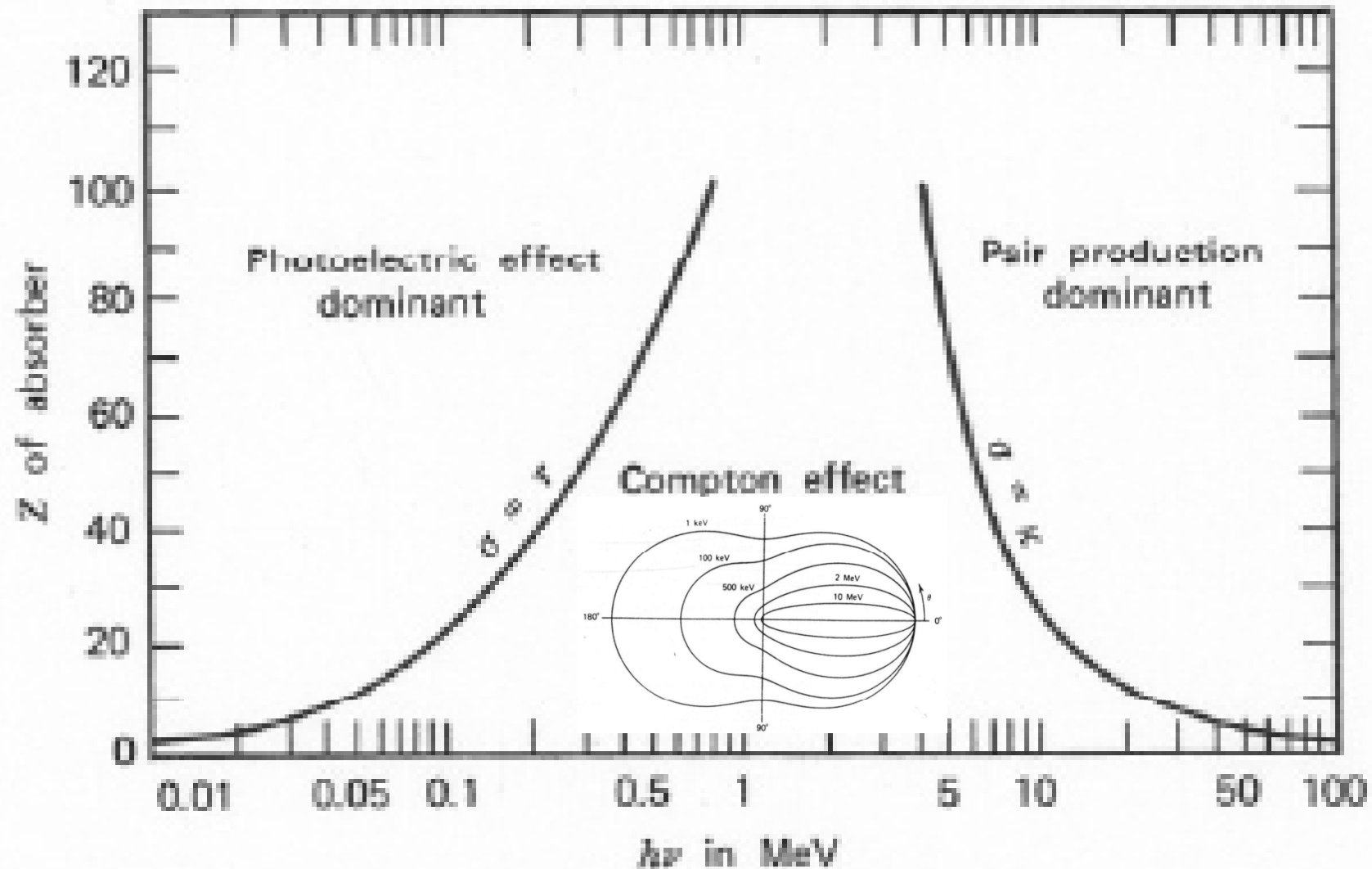
$$I(X) = I_0 e^{-\Sigma_t X}$$

## HW 2 What is coherent scattering.

$$P_{\text{no-interaction}}(X) = e^{-\Sigma_t X}$$

$$P_{\text{interaction}}(X) = 1 - e^{-\Sigma_t X}$$

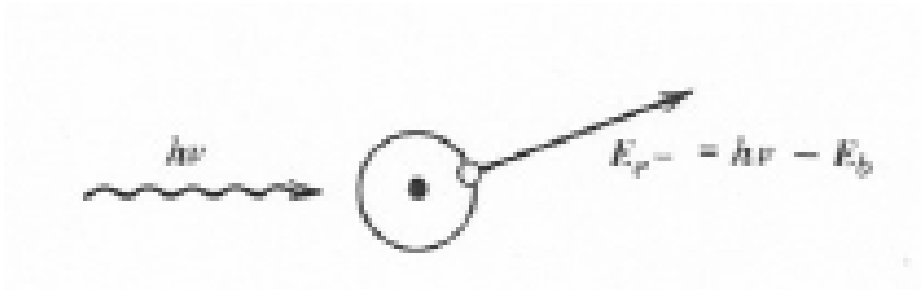
# Principles of Spectrometry



# Principles of Spectrometry

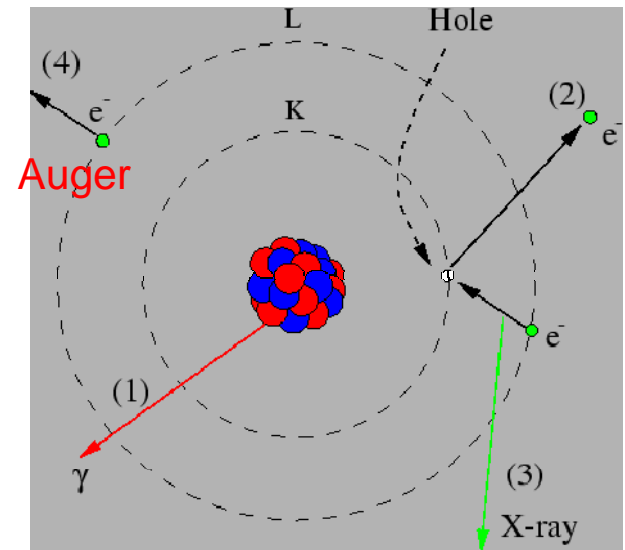
## Photoelectric Absorption.

(with atoms, why??)

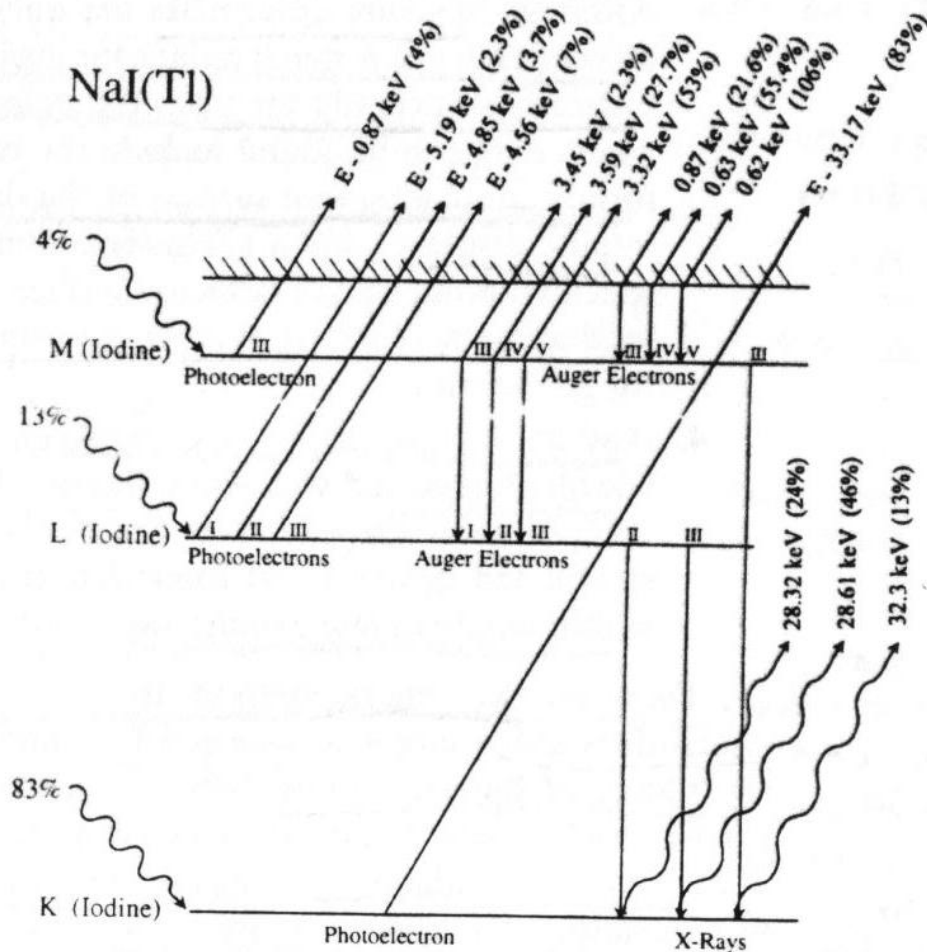


Spectrum if all energy is captured in detector.  
Allows identification of gamma energy.

- Large Detector (in depth?!).
- Photo Peak or Full Energy Peak...?
- Doppler.
- Cost.
- Crystal growth.



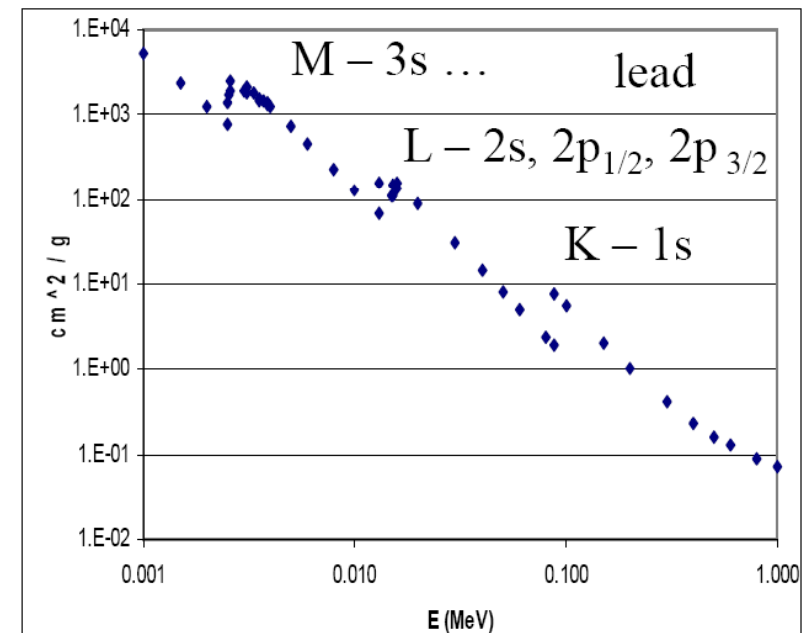
# Principles of Spectrometry



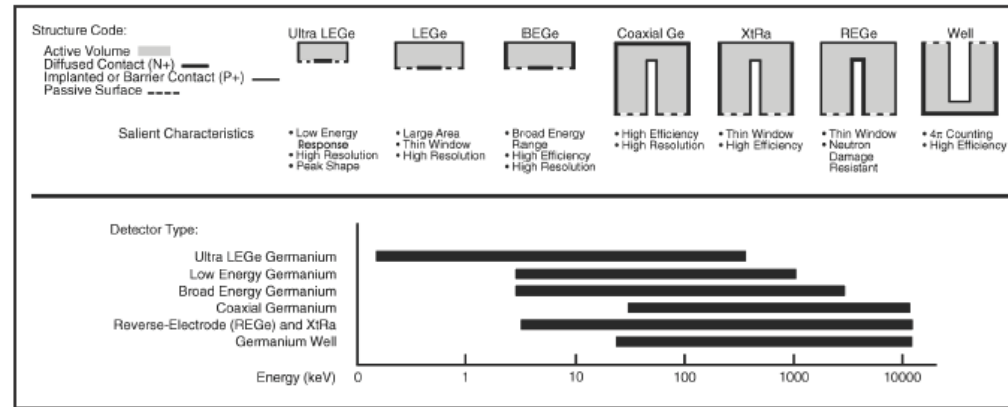
- Spectroscopy, energy deposition.
- Consider what might escape.
- Size and material of detector.
- Shield lining.



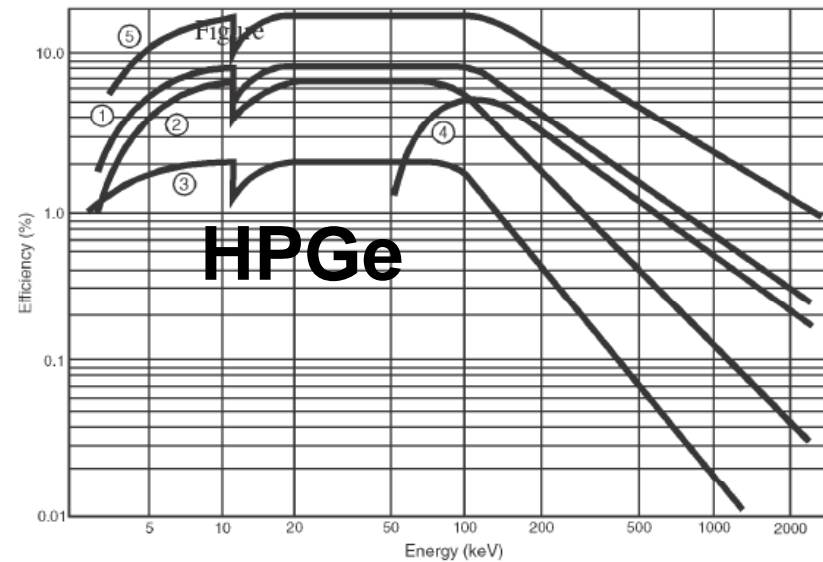
Considerations for  
**shields and detectors?**



# Principles of Spectrometry



## HW 3





# Principles of Spectrometry

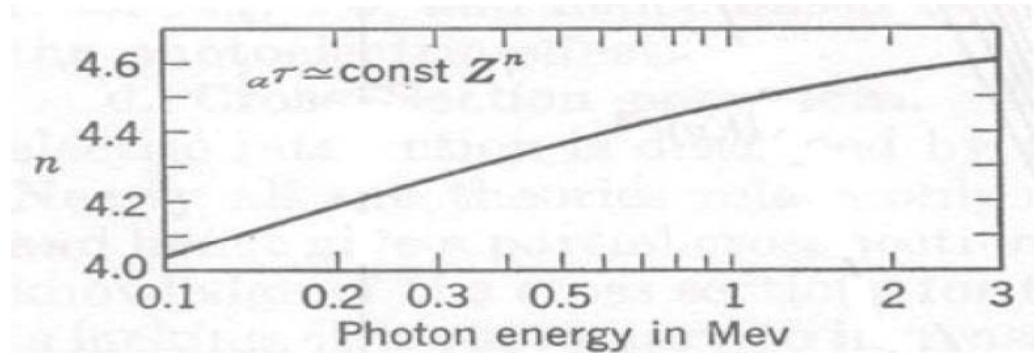
- “Probability” for **photoelectric** absorption:

**Theoretical**

$$\sigma_{\tau} = \int d\Omega \frac{d\sigma_{\tau}}{d\Omega} = 4\sqrt{2}\sigma^0 \frac{Z^5}{(137)^4} \left( \frac{m_e c^2}{\hbar\omega} \right)^{7/2}$$

**Practical**

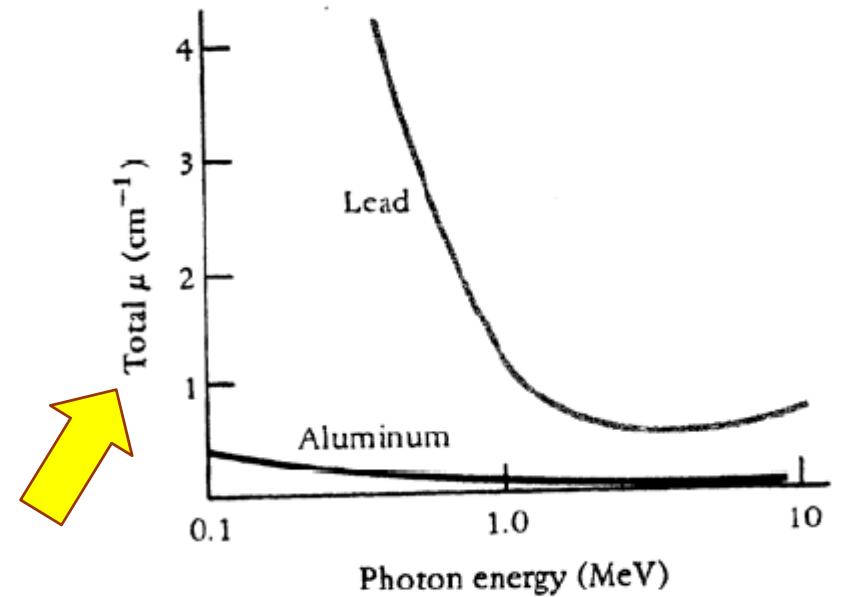
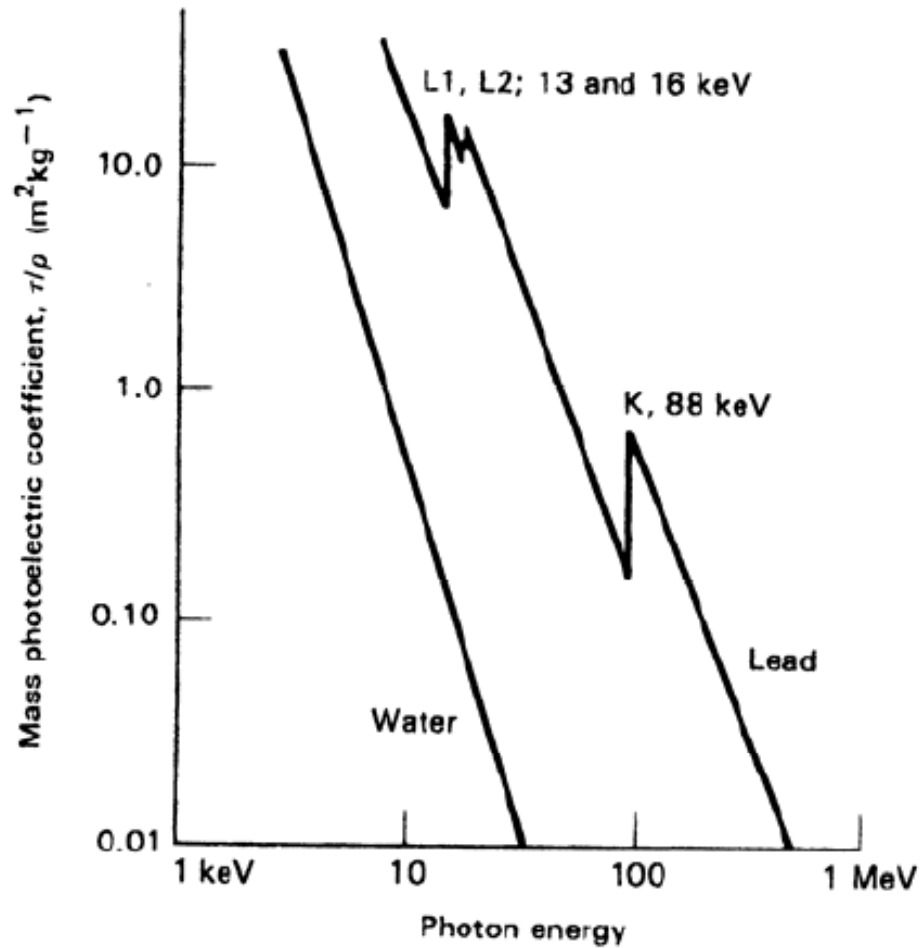
$$\tau \propto \begin{cases} \frac{Z^4}{(h\nu)^{3.5}} & \text{low energy} \\ \frac{Z^5}{(h\nu)^{3.5}} & \text{high energy} \end{cases}$$



- **Strong  $Z$  dependence.**
- Considerations for **shields and detectors.**

# Principles of Spectrometry

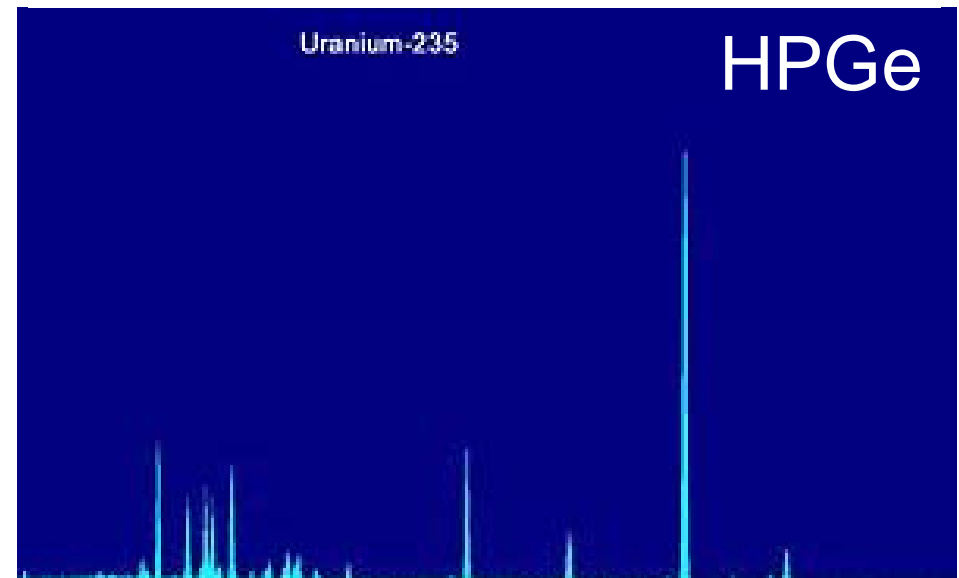
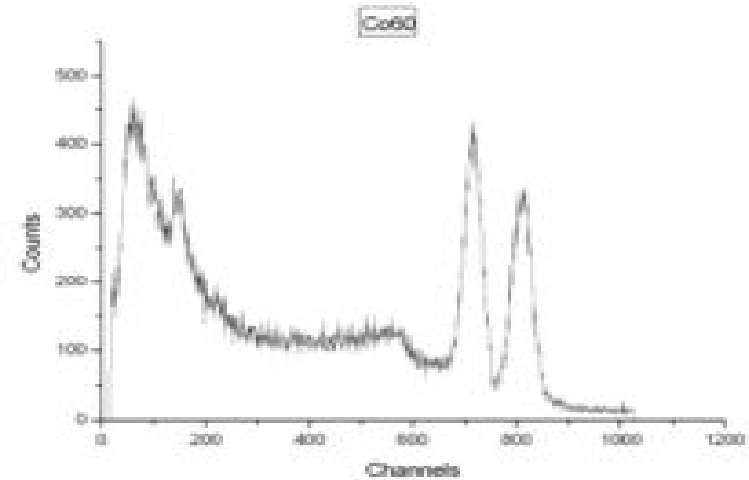
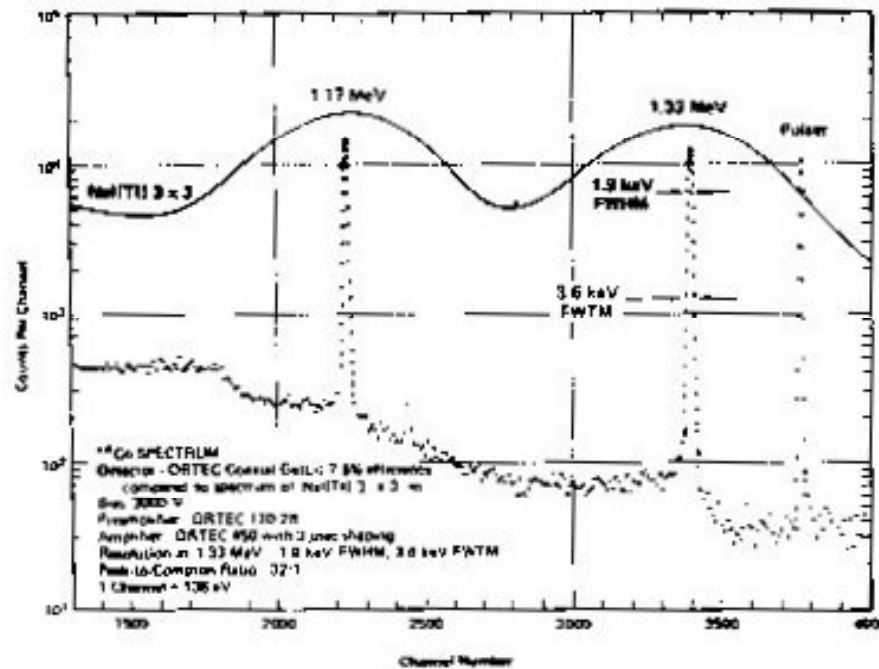
## Efficiency considerations.



# Principles of Spectrometry

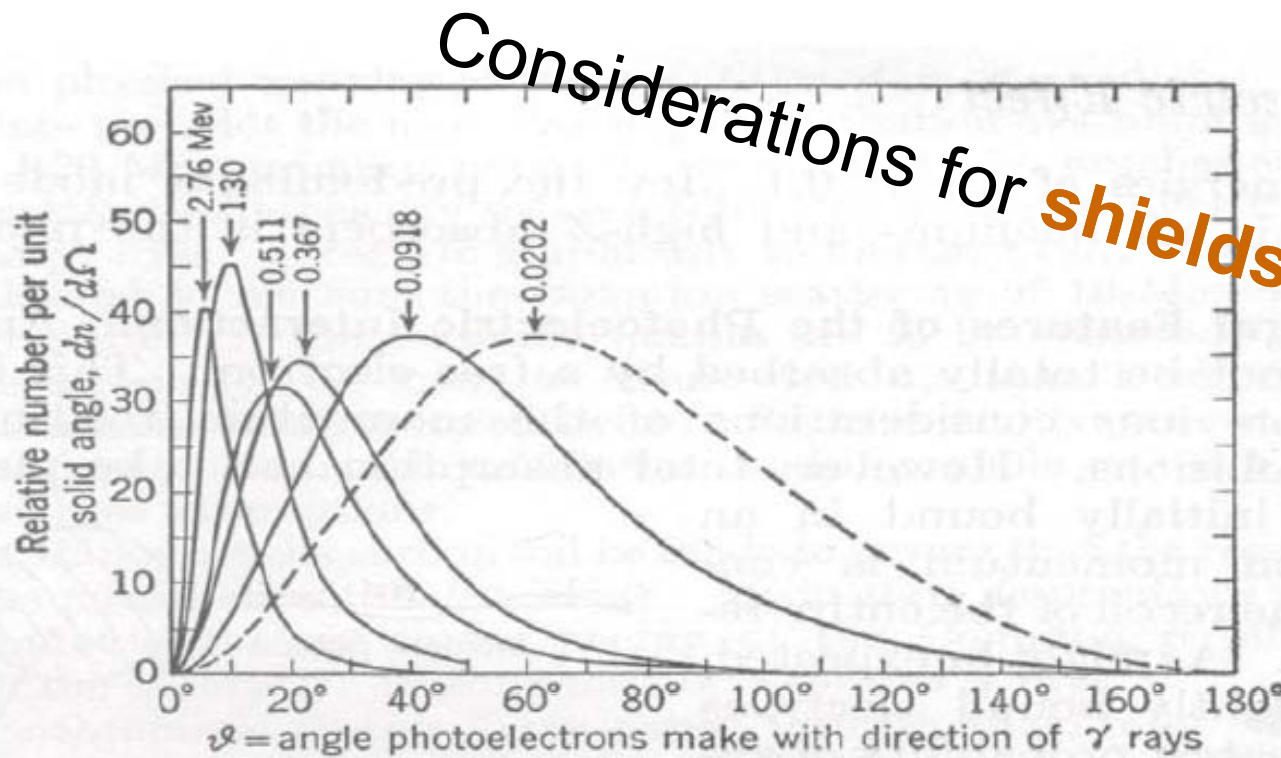
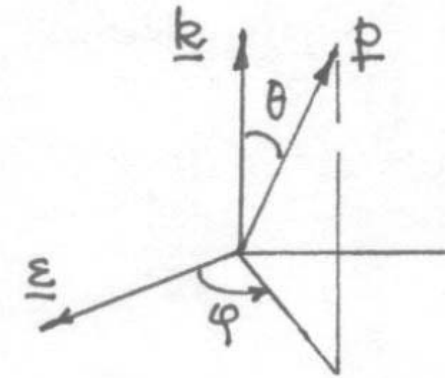
NaI or BGO

## Efficiency vs. Resolution



# Principles of Spectrometry

$$\frac{d\sigma_{\tau}}{d\Omega} = 4\sqrt{2} \frac{r_e^2 Z^5}{(137)^4} \left( \frac{m_e c^2}{\hbar\omega} \right)^{7/2} \frac{\sin^2 \theta \cos^2 \varphi}{\left( 1 - \frac{v}{c} \cos \theta \right)^4}$$



Considerations for **shields and detectors?**